

Hinode/SOT Observations of Prominence Dynamics



Thomas Berger, T. J. Okamoto, T. Magara, A. M. Title,
T. D. Tarbell, S. Tsuneta

and the rest of the SOT Team

Prominences Dynamics

Font size indicates relative novelty

1. Downflow streams
2. Vortex flows
3. Large-scale oscillations (distinct from Alfvénic oscillations in AR filaments)
4. Counterstreaming flows
5. Large-scale “bubbles”
6. Turbulent upflow plumes

1. Downflow streams
2. Vortex flows
3. Large scale oscillations
4. Counterstreaming flows
5. Large scale bubble
6. Turbulent upflows

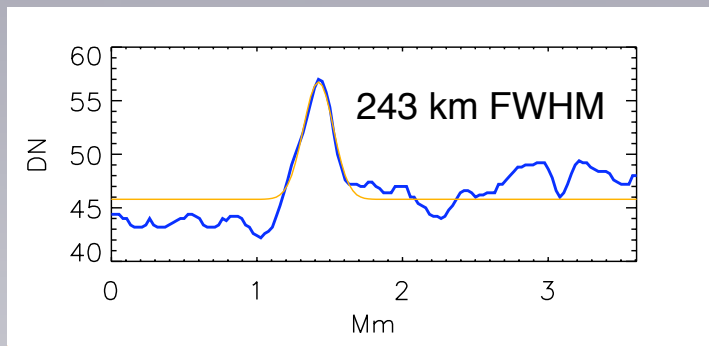
Ca II H-line 396.8nm
30-Nov-2006
NW limb 6 hrs.

30-Nov-2006
01:10:31 UT



Downflow Stream

Grid = 2"

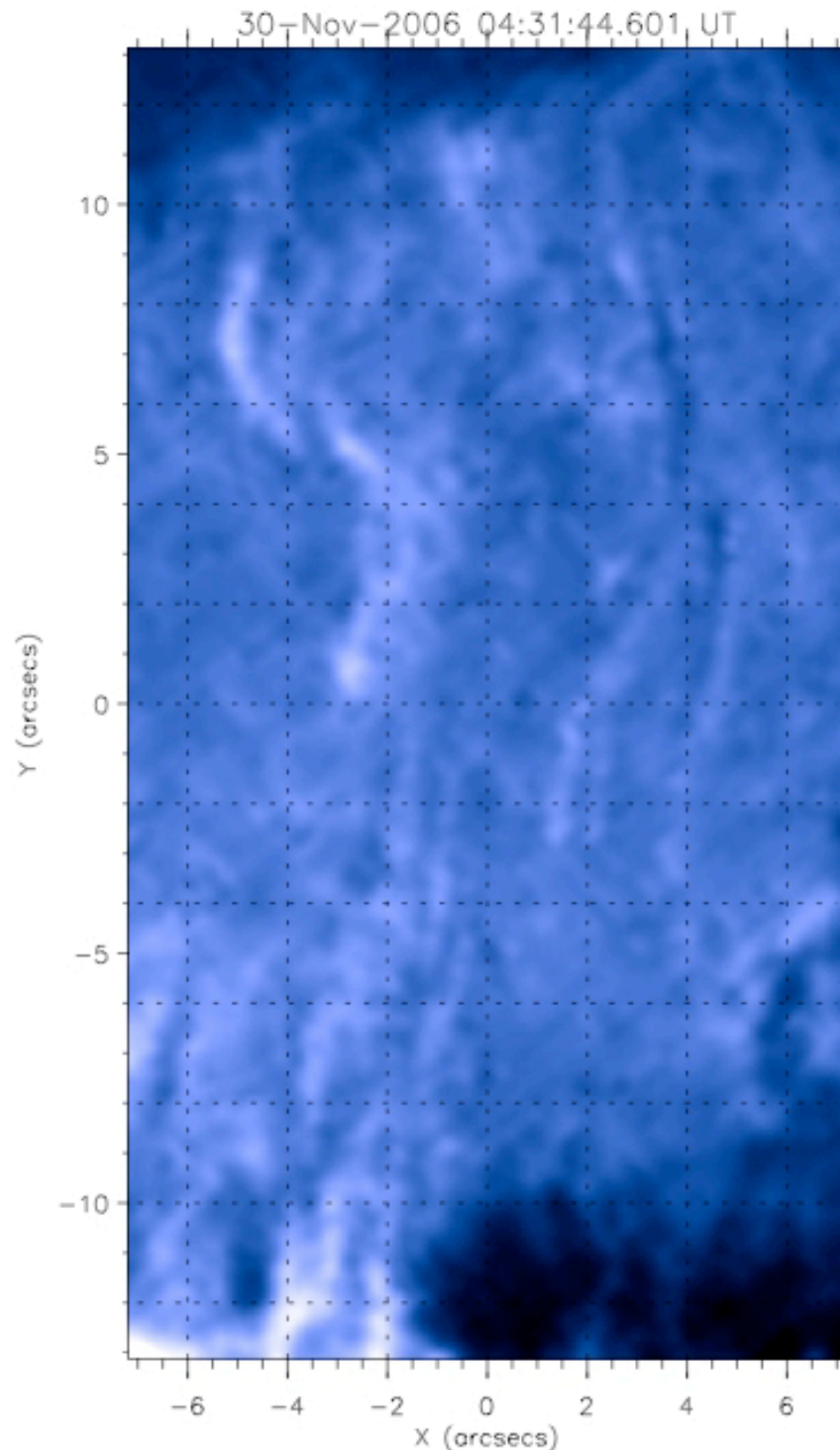


- Speed $\sim 10 \text{ km s}^{-1}$
- Distance $\sim 20\text{--}30 \text{ Mm}$
- Duration $\sim 10 \text{ min}$
- Width $\sim 200\text{--}700 \text{ Mm}$

Ubiquitous in most QPs

Note: $V_{\text{freefall}(20 \text{ Mm})} \sim 100 \text{ km s}^{-1}$.
 $V \ll V_{\text{freefall}}$, like waterfalls...

Theories that neglect drag force are not applicable to this problem.



Vortex Flow

Grid = 2"

2.5 Rotations

Rate = 3.27×10^{-3} rad/sec

~3000 km diameter

Assume:

$T = 7000$ K

$n_e = 10^{11}$ cm⁻³

$f = 0.3$

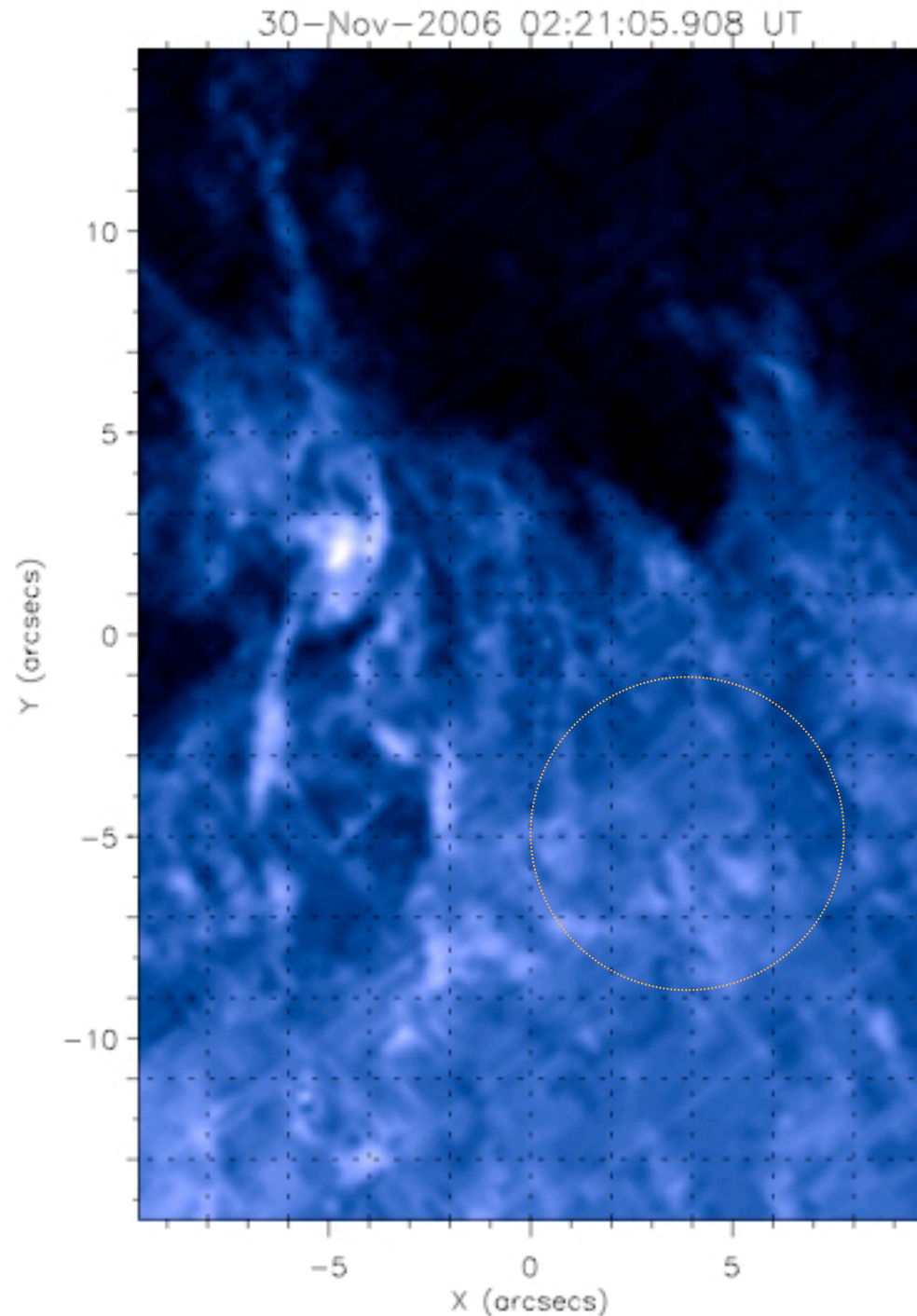
$B = 10$ G

Then:

$P = 0.042$ Pa

$\beta \sim 0.1$

Must explain vortex motion
in low- β environment. Are the
field lines doing this?



Upflow Plumes

Grid = 2"

Speed $\sim 20 \text{ km s}^{-1}$

Distance $\sim 10\text{--}20 \text{ Mm}$

Duration $\sim 10 \text{ min}$

Width $\sim 400\text{--}2000 \text{ Mm}$

Intermittent

Assume:

$T = 7000 \text{ K}$

$n_e = 10^{11} \text{ cm}^{-3}$

$f = 0.3$

$B = 10 \text{ G}$

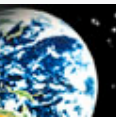
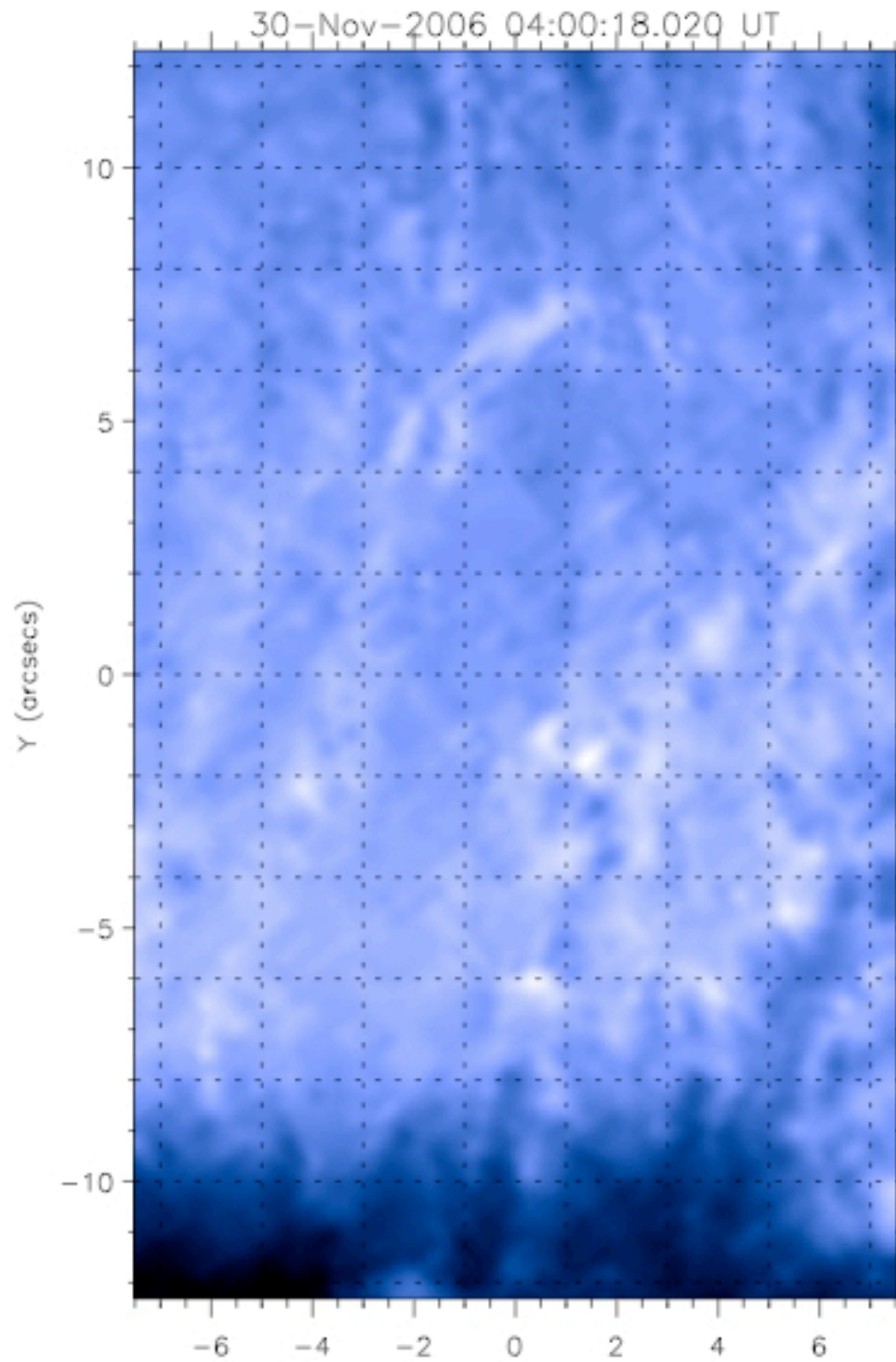
Then:

$Re \sim 400,000$

but...

$P = 0.042 \text{ Pa}$

$\beta \sim 0.1$



Upflow Plumes

Grid = 2"

Speed $\sim 20 \text{ km s}^{-1}$

Distance $\sim 10\text{--}20 \text{ Mm}$

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Intermittent

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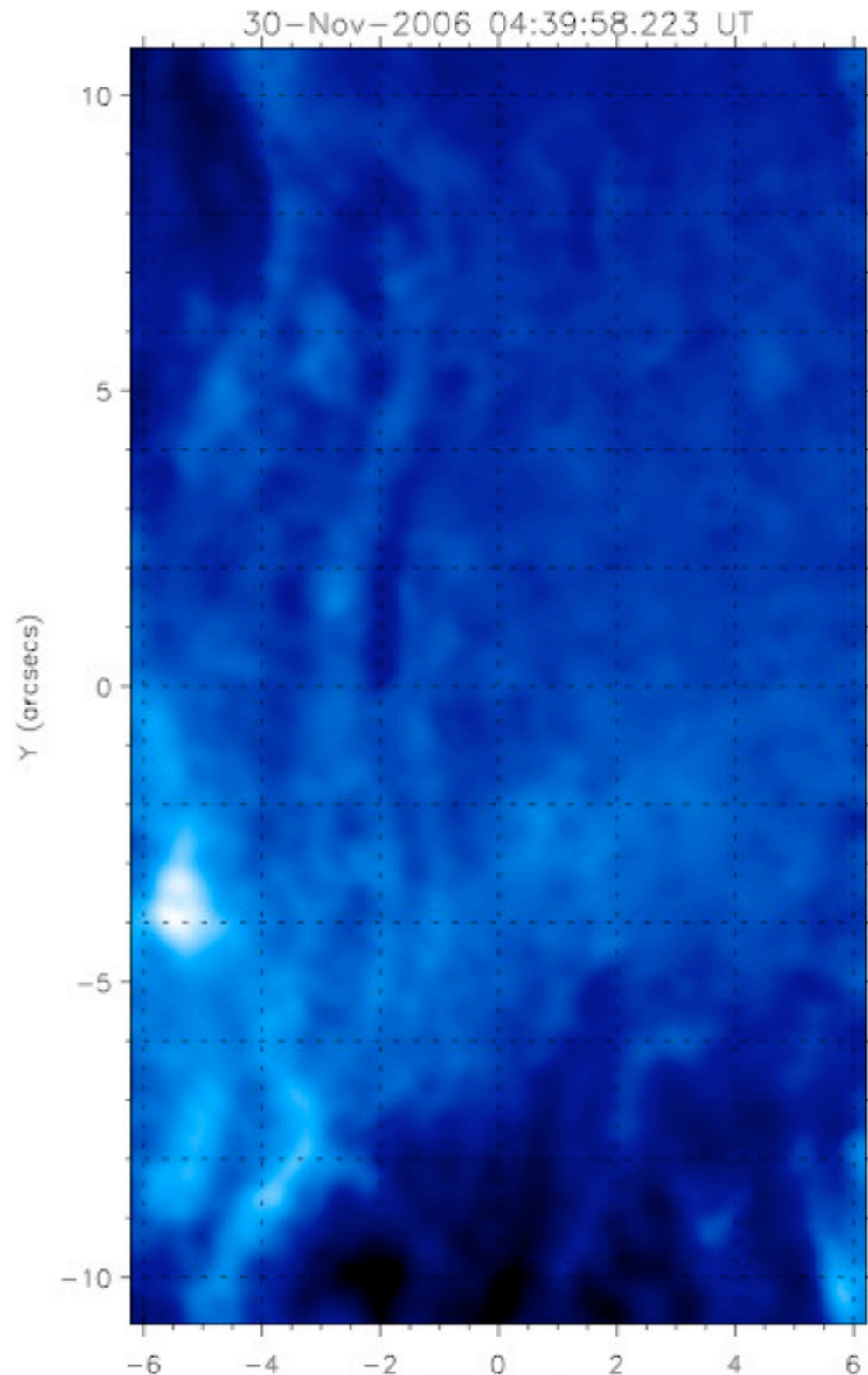
Then:

$Re \sim 400,000$

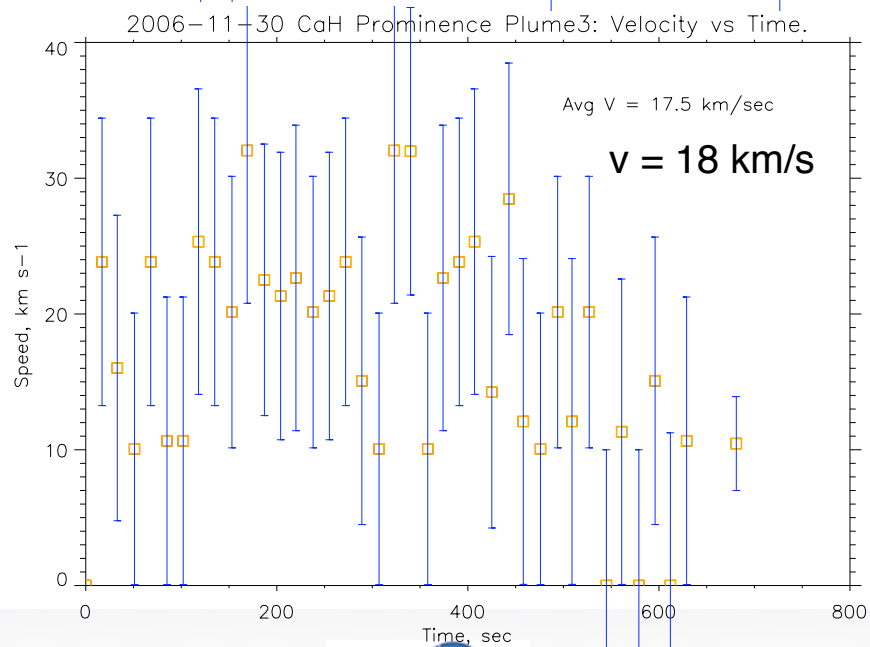
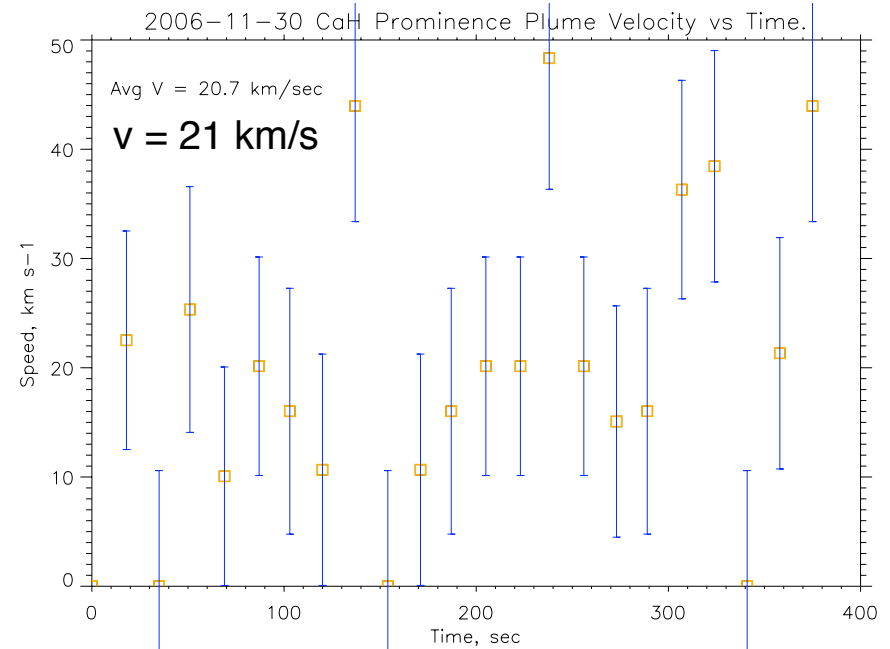
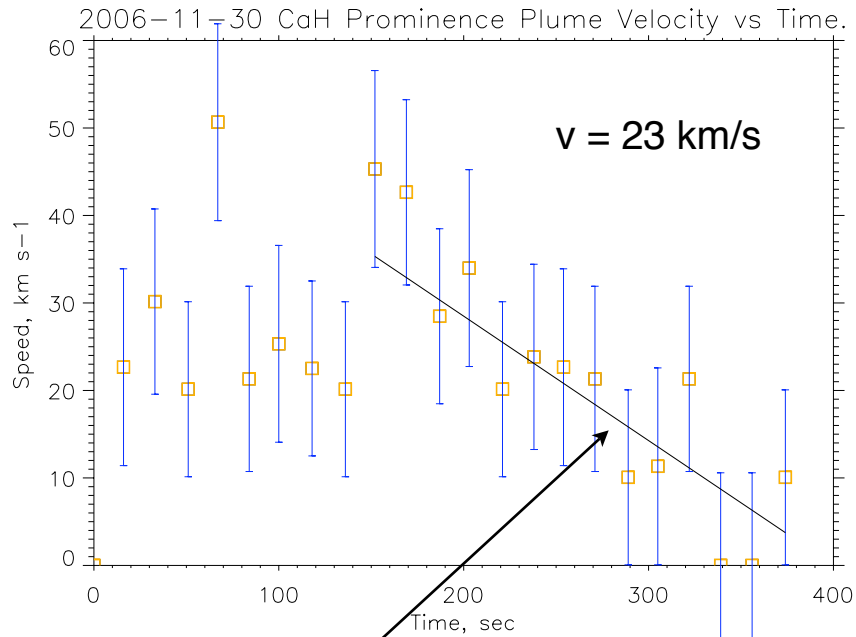
but...

$P = 0.042 \text{ Pa}$

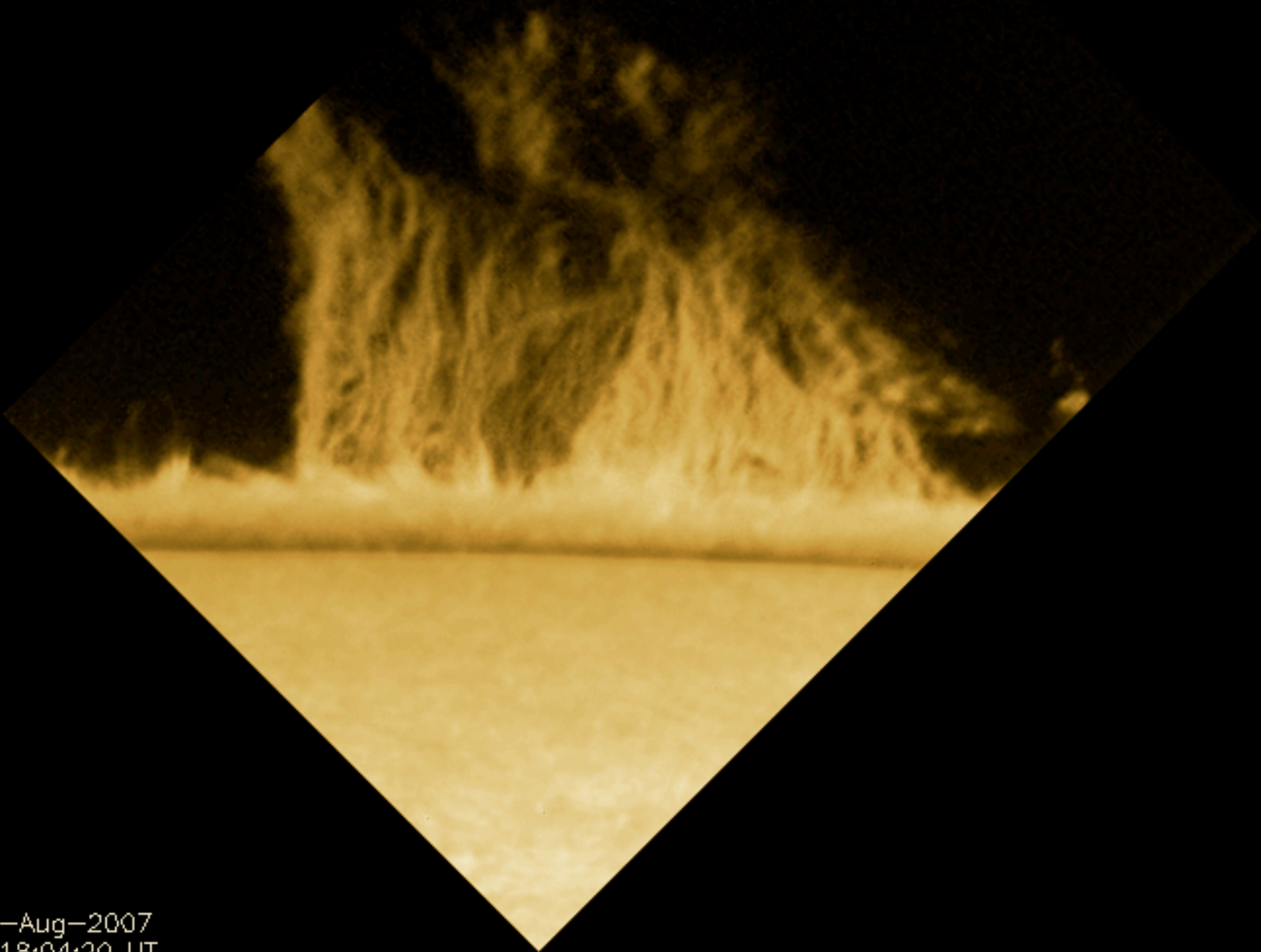
$\beta \sim 0.1$



Upflow Plume Velocity Measurements



6. Turbulent Upflow Plumes

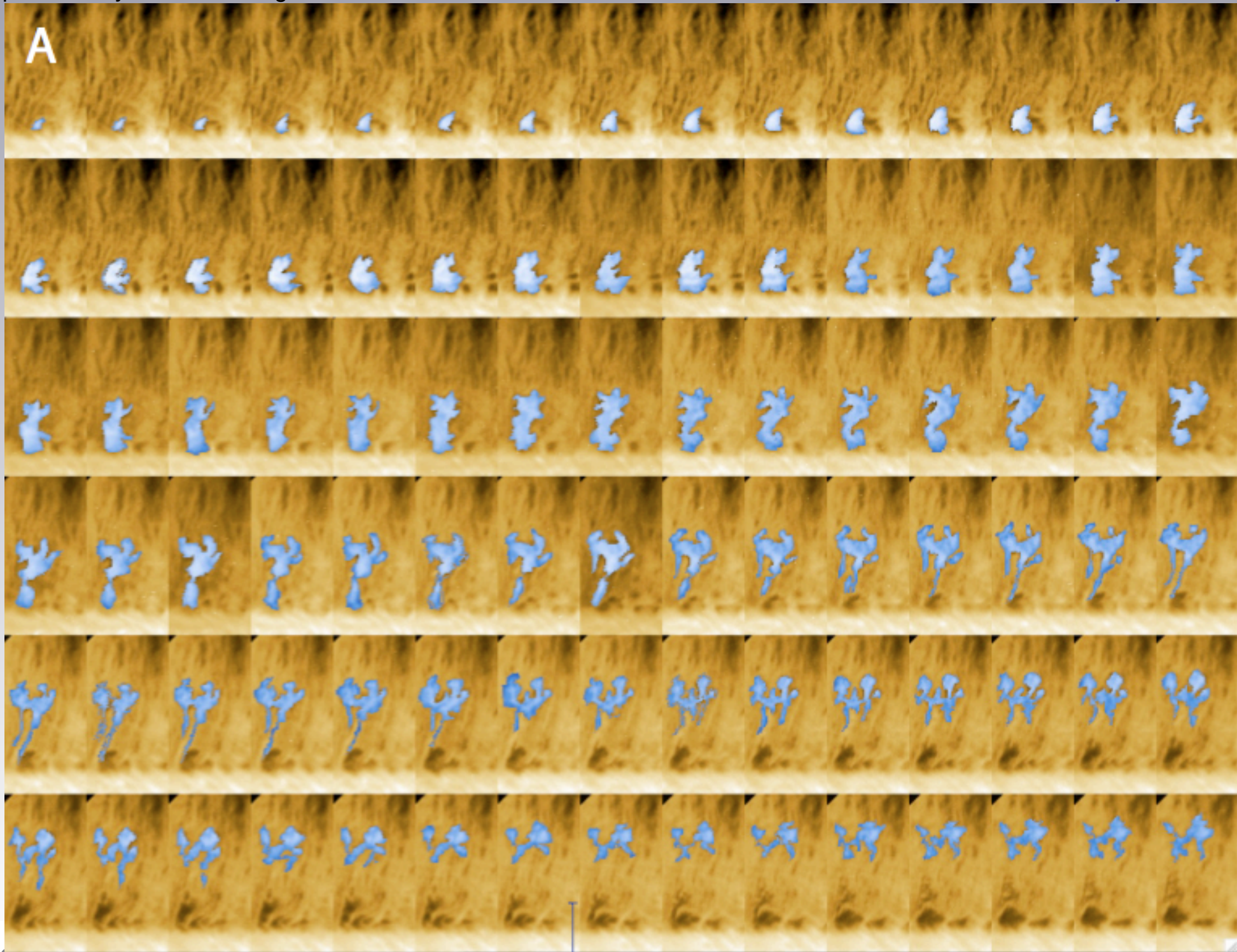


8-Aug-2007
18:04:20 UT



H-alpha 656.3nm
8-Aug-2007
NE limb 4 hrs.

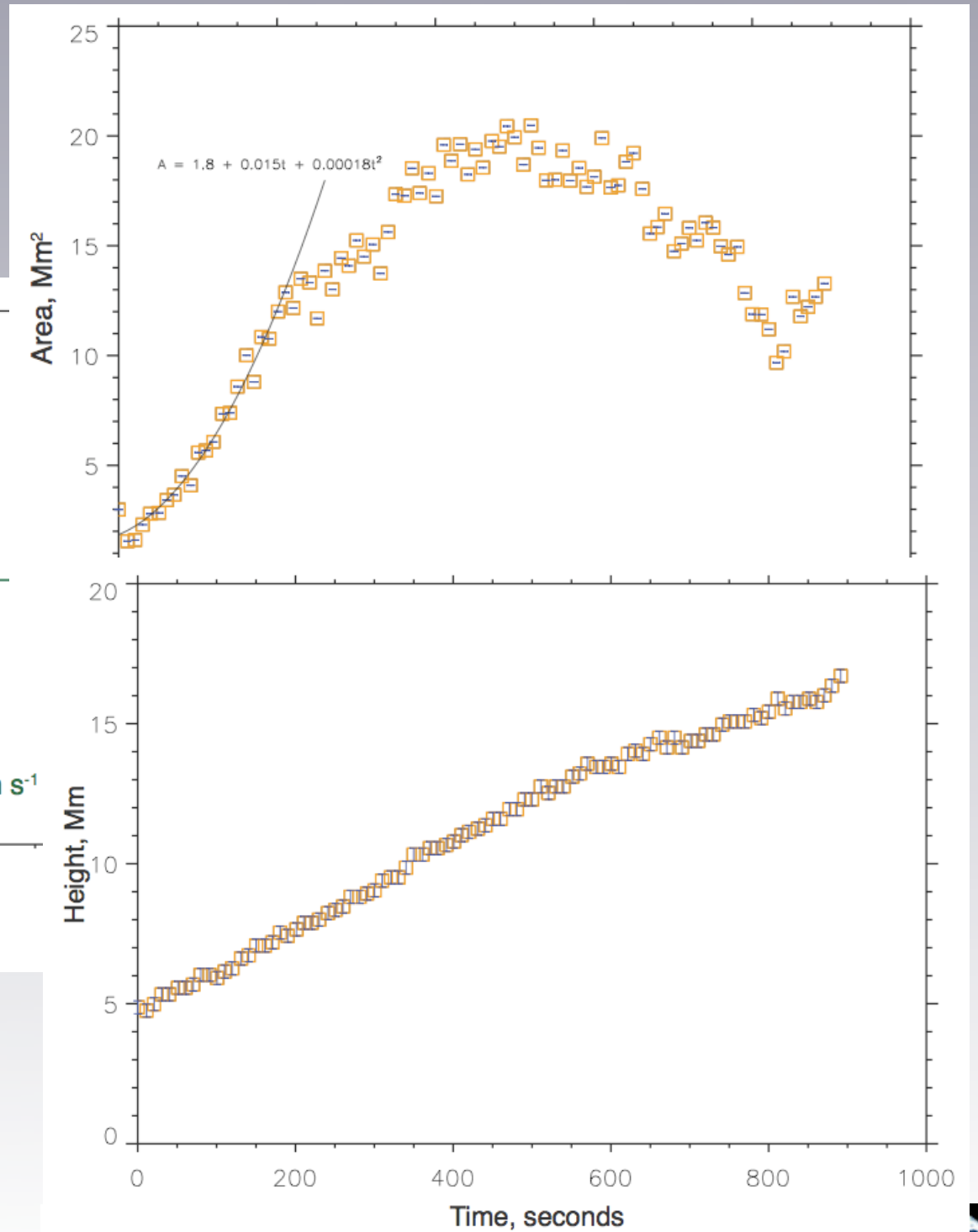
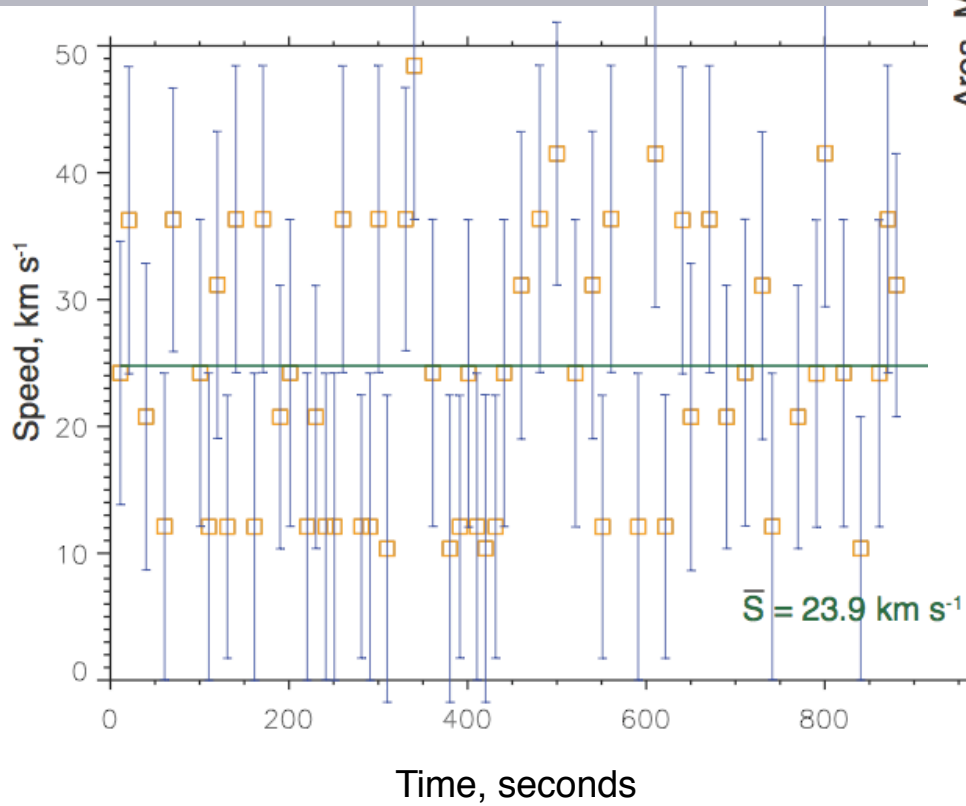


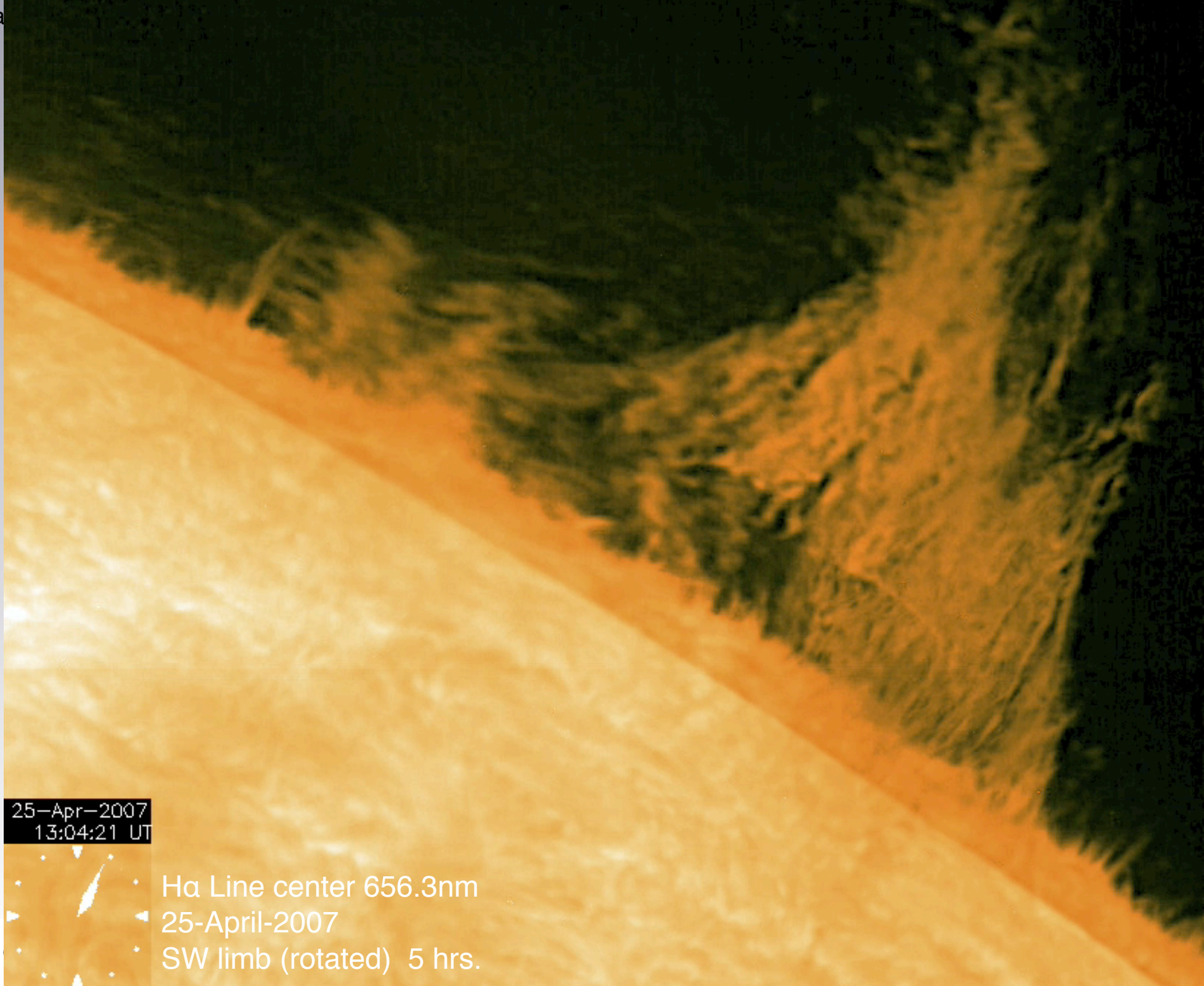


LOC



Upflow Plume Growth Rate





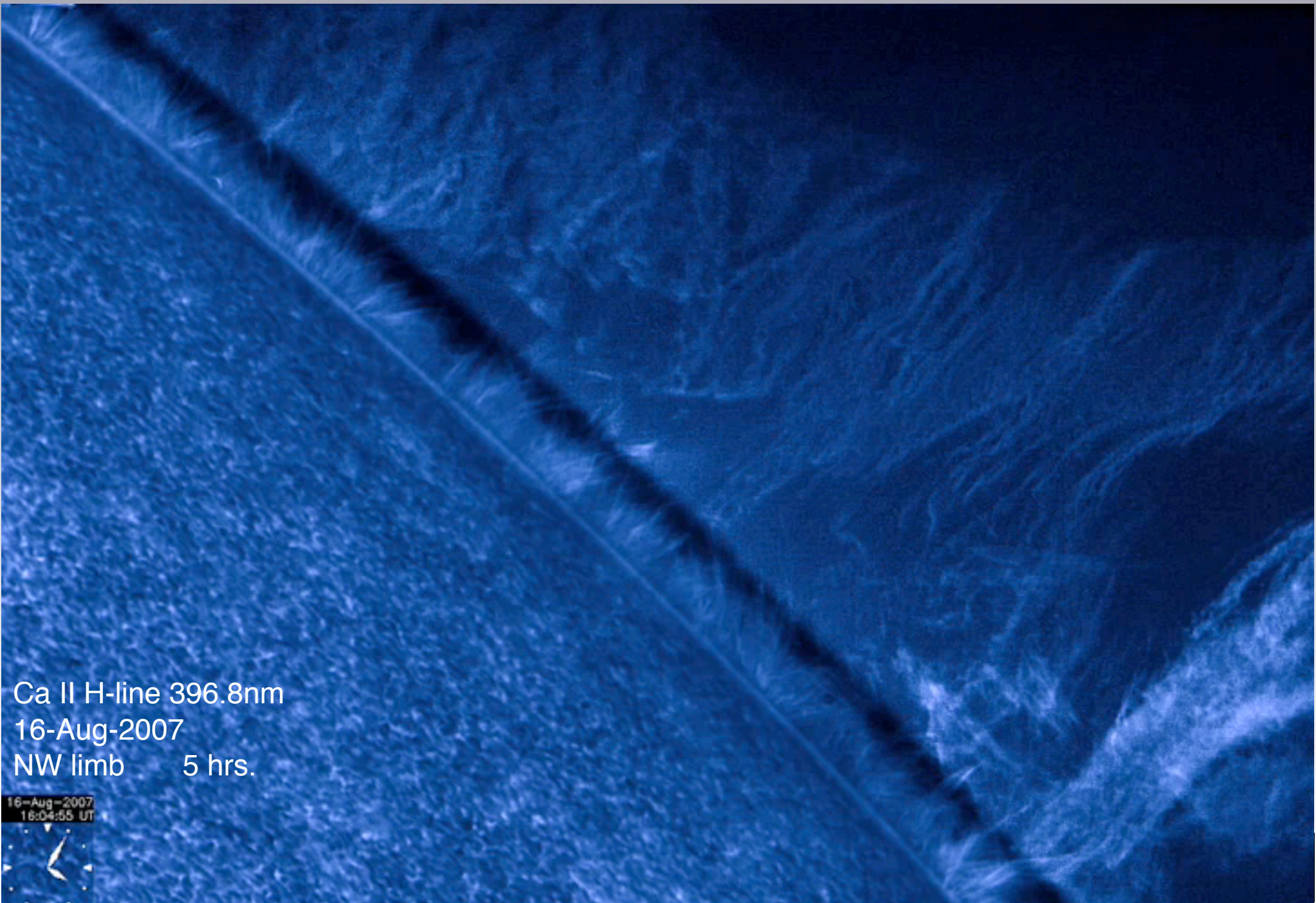
25-Apr-2007
13:04:21 UT



Ha Line center 656.3nm
25-April-2007
SW limb (rotated) 5 hrs.



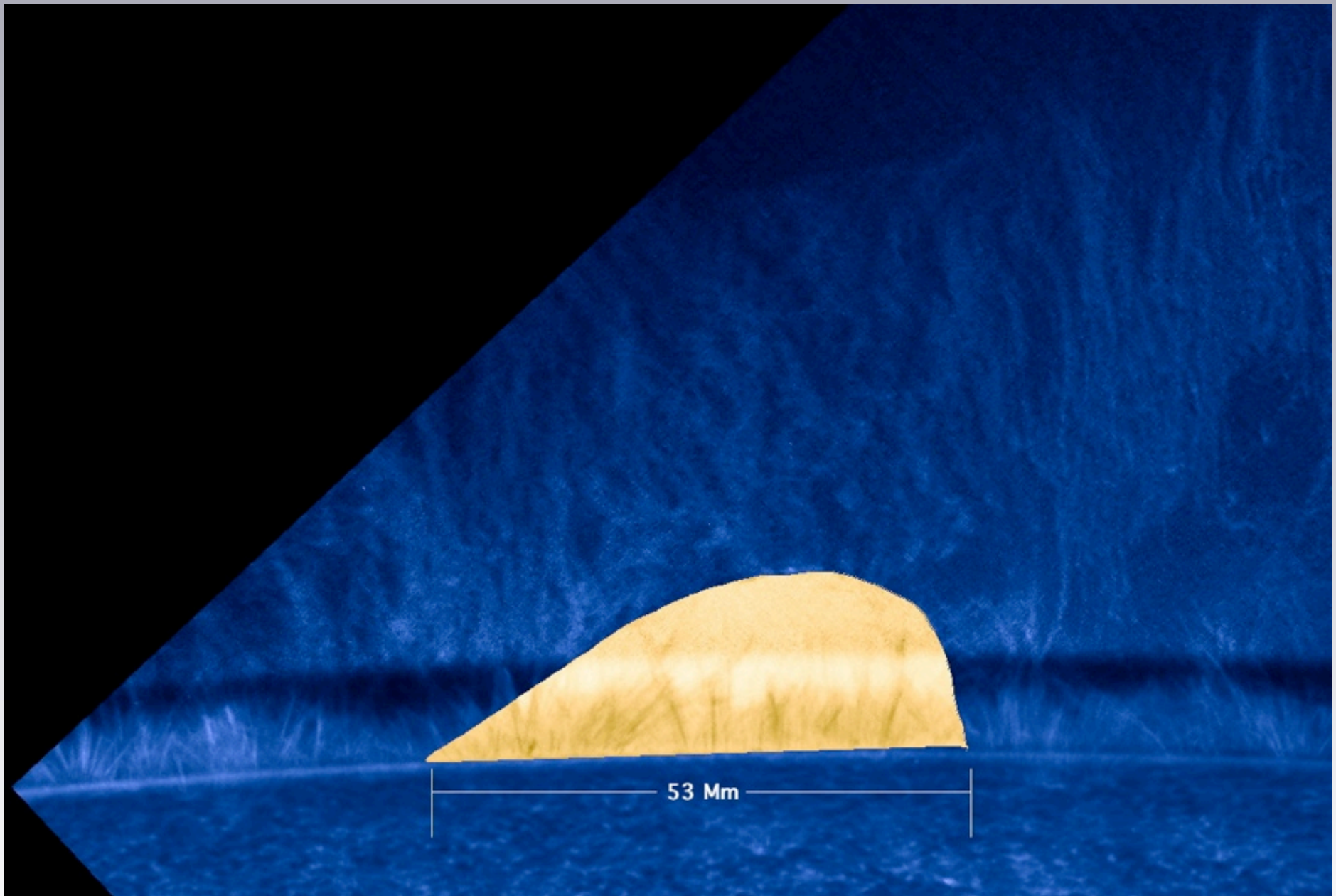
5. Large-scale Bubbles



Ca II H-line 396.8nm
16-Aug-2007
NW limb 5 hrs.

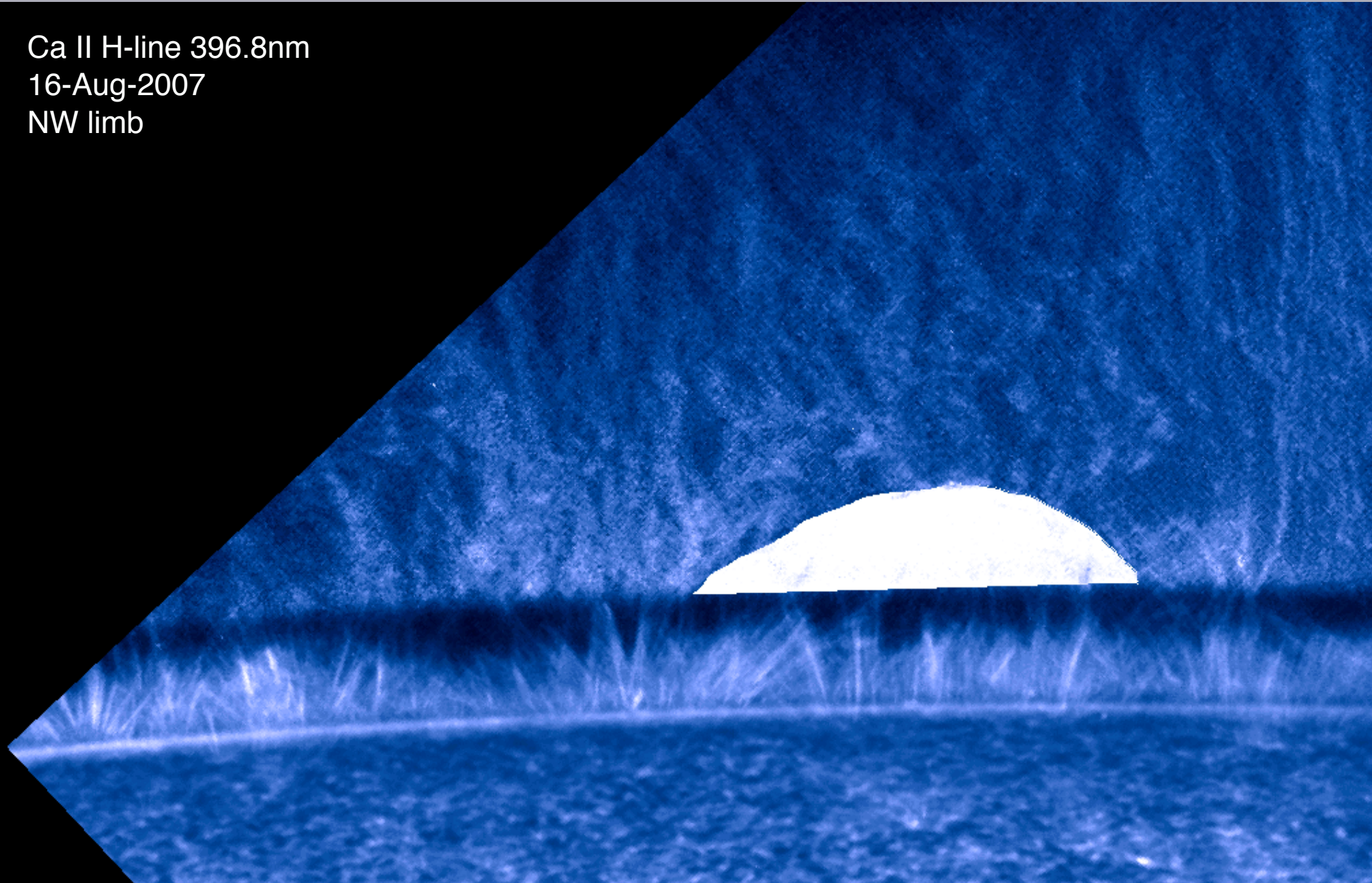


5. Large-scale Bubbles

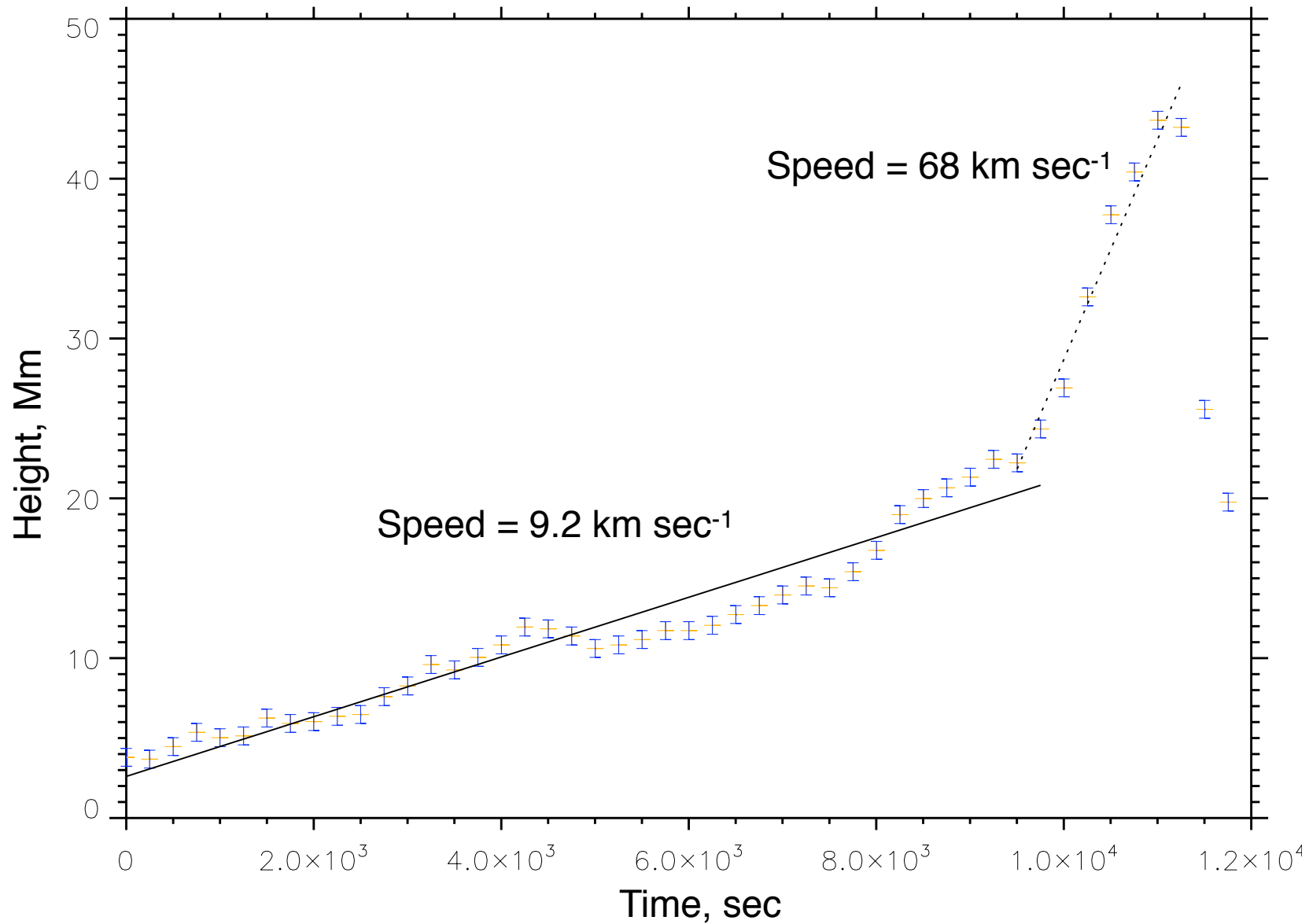


5. Large-scale Bubbles

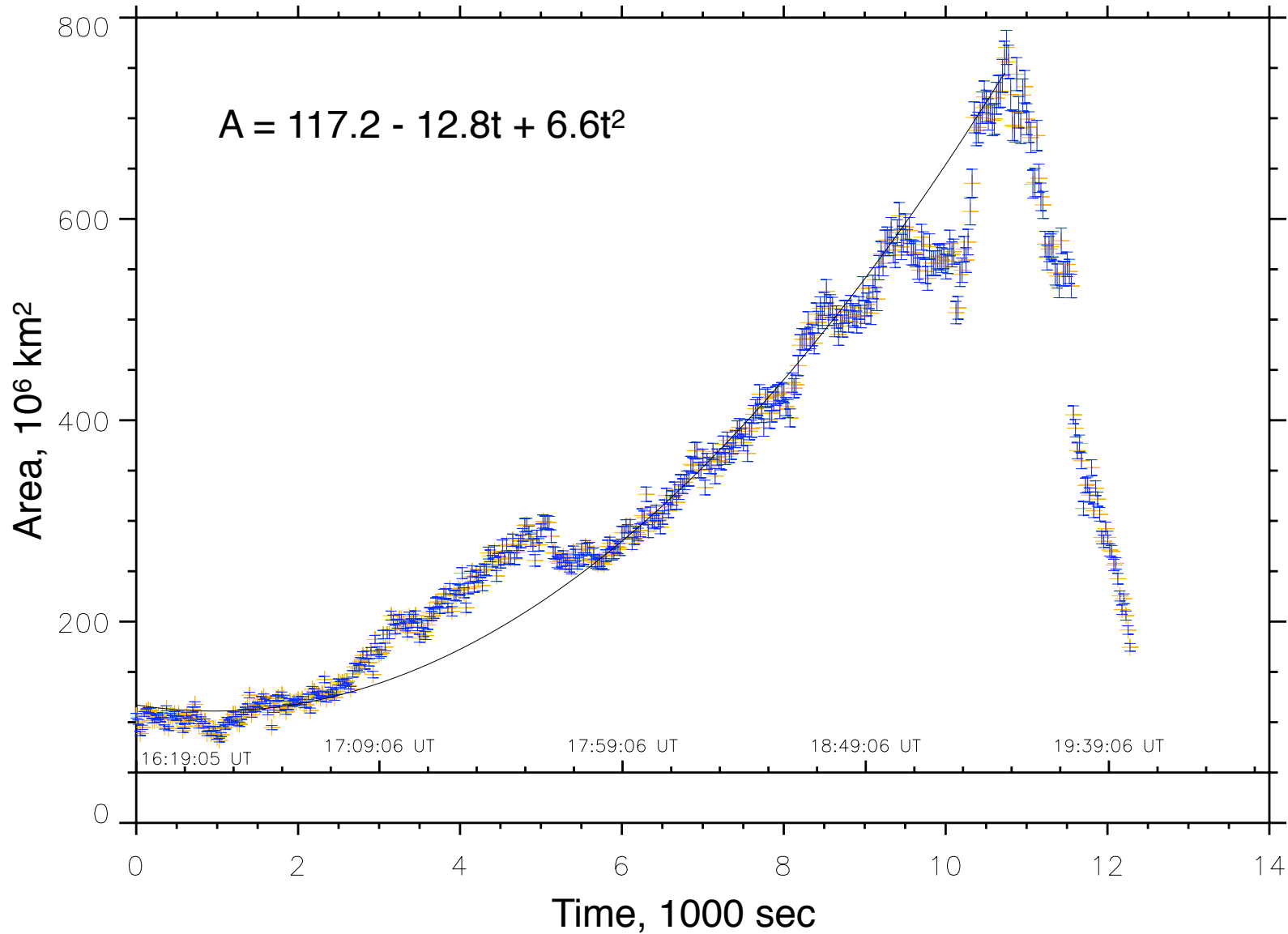
Ca II H-line 396.8nm
16-Aug-2007
NW limb



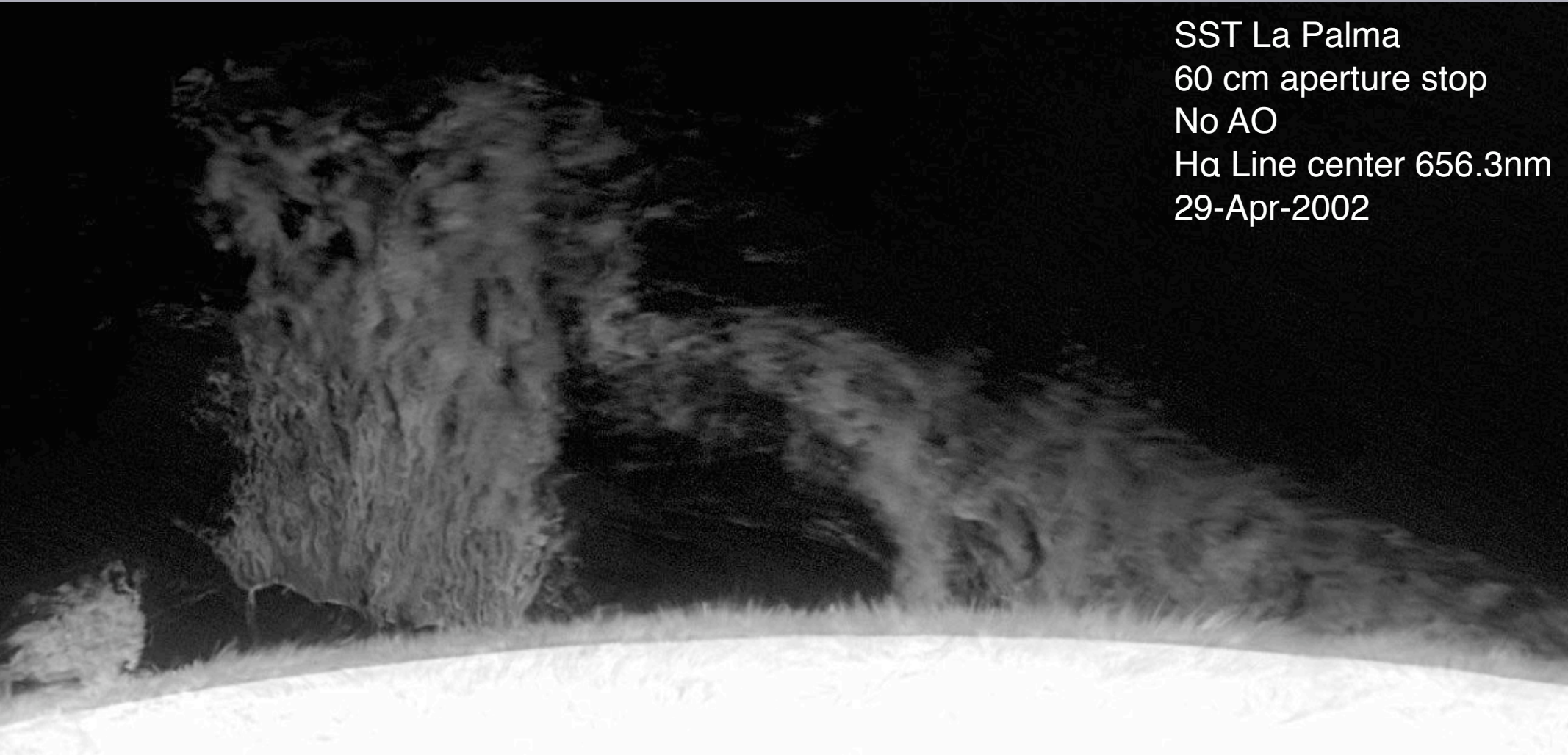
Large Bubble Growth Rate



Large Bubble Growth Rate



Why we need optical instruments in space AND on the ground



SST La Palma
60 cm aperture stop
No AO
H α Line center 656.3nm
29-Apr-2002

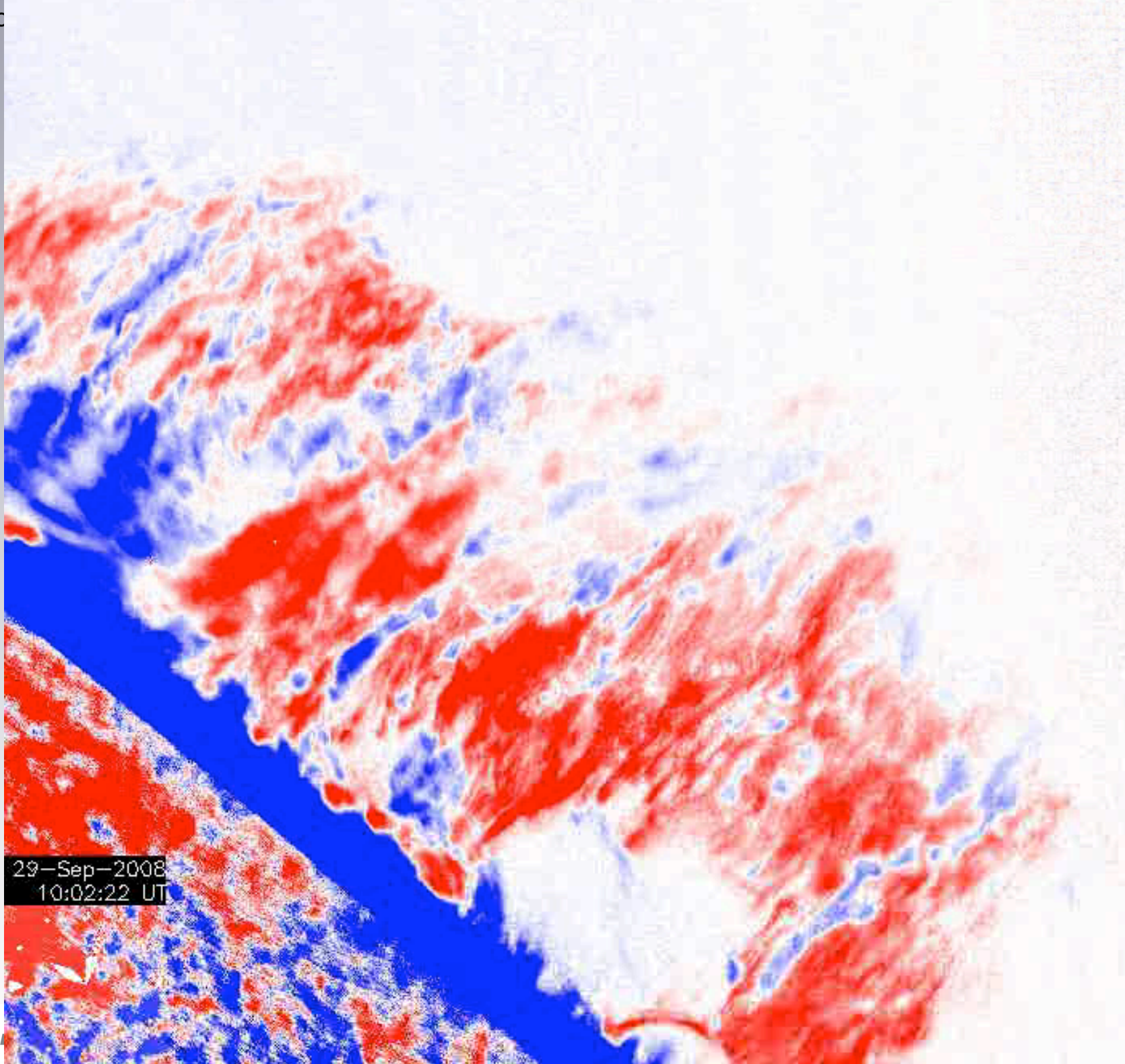
Much larger FOV, comparable resolution

Ha wing 656.31nm
29-Sep-2008
NW limb 4 hrs.
30 sec cadence

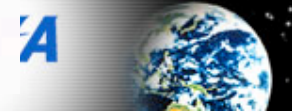


29-Sep-2008
10:02:02 UT



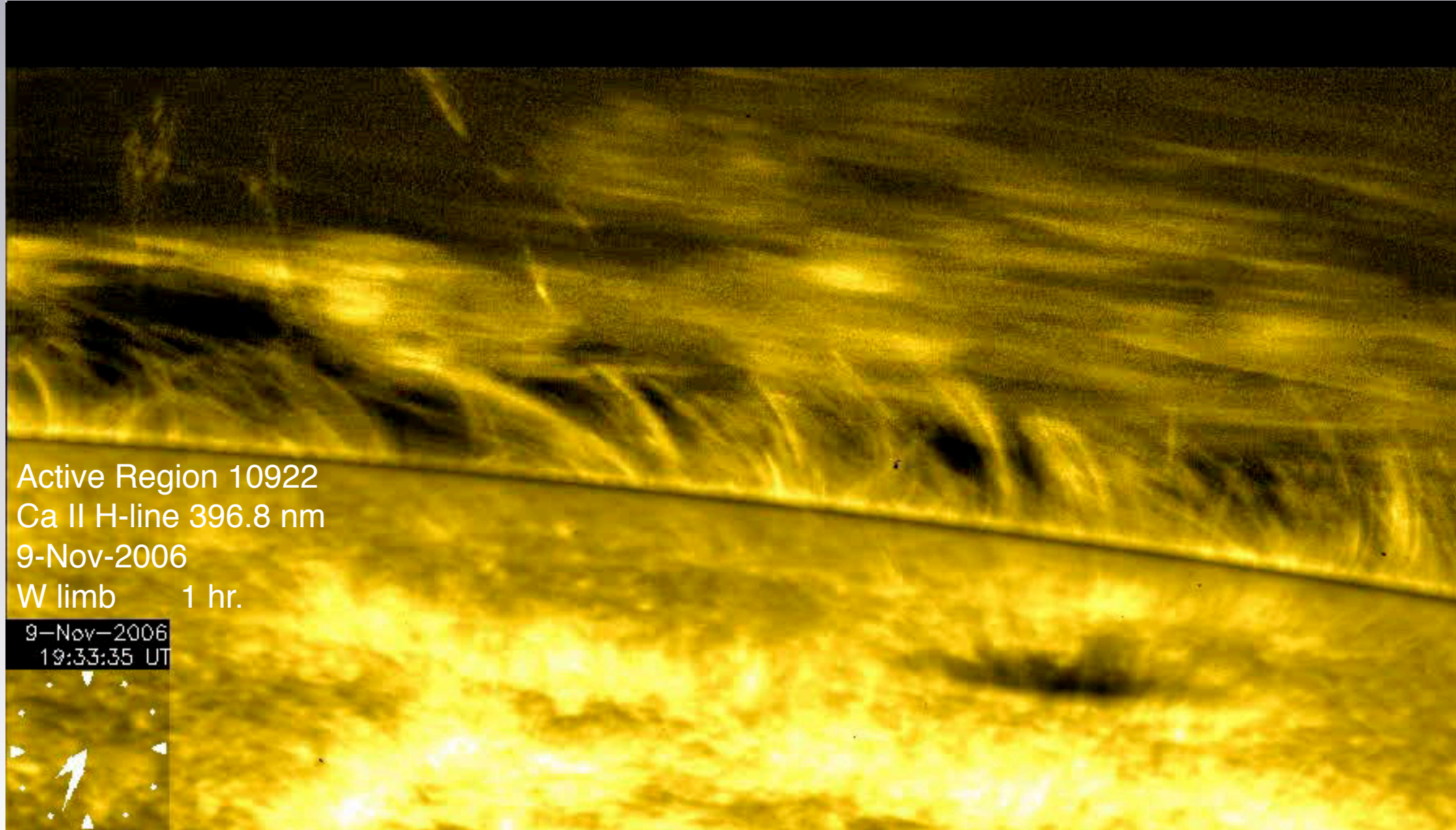


29-Sep-2008
10:02:22 UT



Active Region Filaments

primarily counterstreaming on horizontal filaments



Active Region 10922
Ca II H-line 396.8 nm
9-Nov-2006
W limb 1 hr.

9-Nov-2006
19:33:35 UT



Active Region Filaments

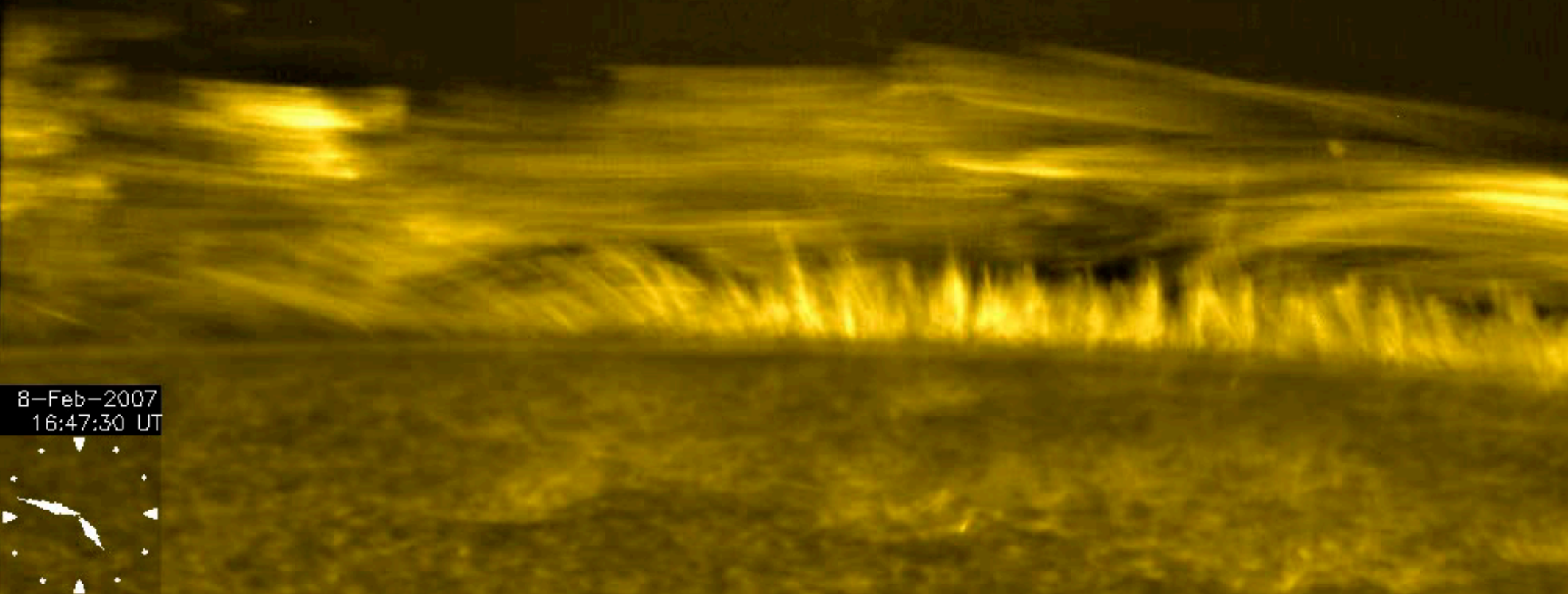
primarily counterstreaming on horizontal filaments

Active Region 10940

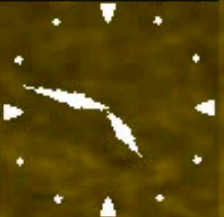
Ca II H-line 396.8 nm

08-Feb-2007

W limb ~ 2 hrs.



8-Feb-2007
16:47:30 UT



Conclusions

Quiescent Prominences always show one or more of the 6 dynamics.

- Bright downflow “streams”, $\langle V \rangle \sim 10$ km/sec, 10 min characteristic lifetime, 250 - 700 km width.
- Vortices, characteristic scale 10^3 km, 3×10^{-3} rad/sec
- Large scale bubble events “inflate” at ~ 10 km/sec. Bursting can be > 50 km/sec.
- Dark upflow “plumes” are intermittent, $\langle V \rangle \sim 20$ km/sec, 10 minute characteristic lifetime, 400 — 2000 km width.

There is no such thing as a static prominence.

There is no need to “suspend gas against gravity”

Active Region prominences are **always** are horizontally structured.

Very little or no vertical motions - “vertically static” - counterstreaming is dominant dynamic.

- No plume or bubble formation.
- Alfvén wave propagation on horizontal filaments.
- Evidence of rotation of groups of filaments.