

Magnetic Implosions and their Consequences

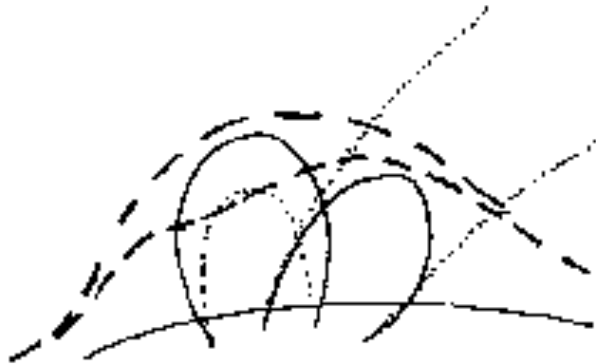
Hugh Hudson

SSL, UC Berkeley and U. Of Glasgow

*The conjecture: a flare or CME must reduce $\int B^2 dV$
in the coronal half-space. Energy is conserved.*

*The significance: the geometry and timing of the
implosion are observables that can guide us to
the nature of the instability.*

Key cartoons

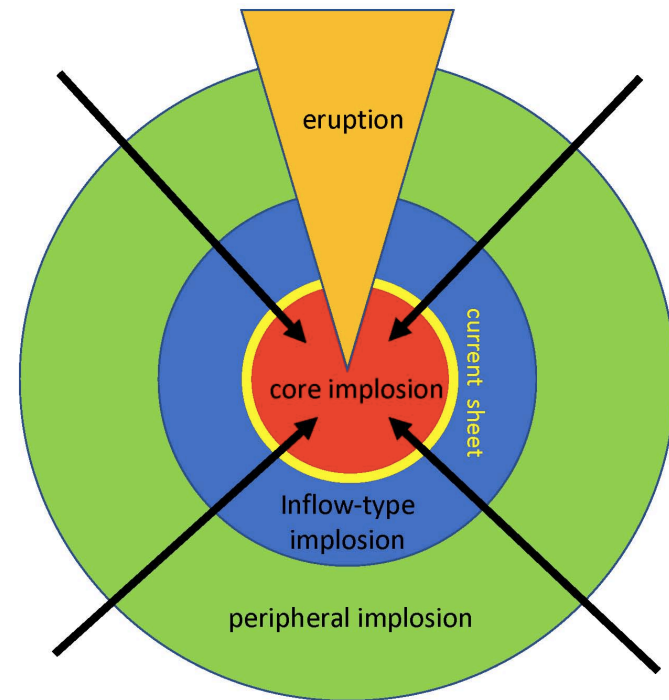


Hudson & Cliver 2001

- - - Magnetoisobar

..... Magnetic field post

___ Magnetic field prior



Wang Juntao 2018

Theoretical history

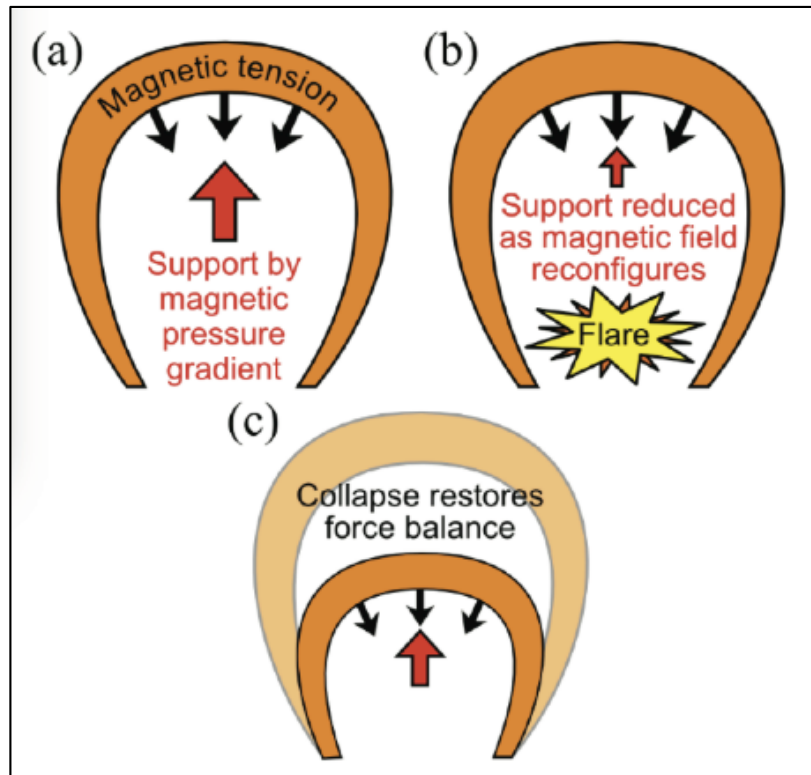
ADS search: abstracts (implosion, flare)

- Sakai, J. [1990ESASP.311....7S](#)
- Craig, I. [1991PASAu...9..225C](#)
- McClymont & Craig [1996ApJ...466..487M](#)
- Hudson, H. [2000ApJ...531L..75H](#)
- Spicer, D., 2000 **“So what? That’s in all the plasma physics textbooks!”**
- Janse & Low [2007A&A...472..957J](#)
- ...
- Barczynski et al. 2019arXiv190405447B

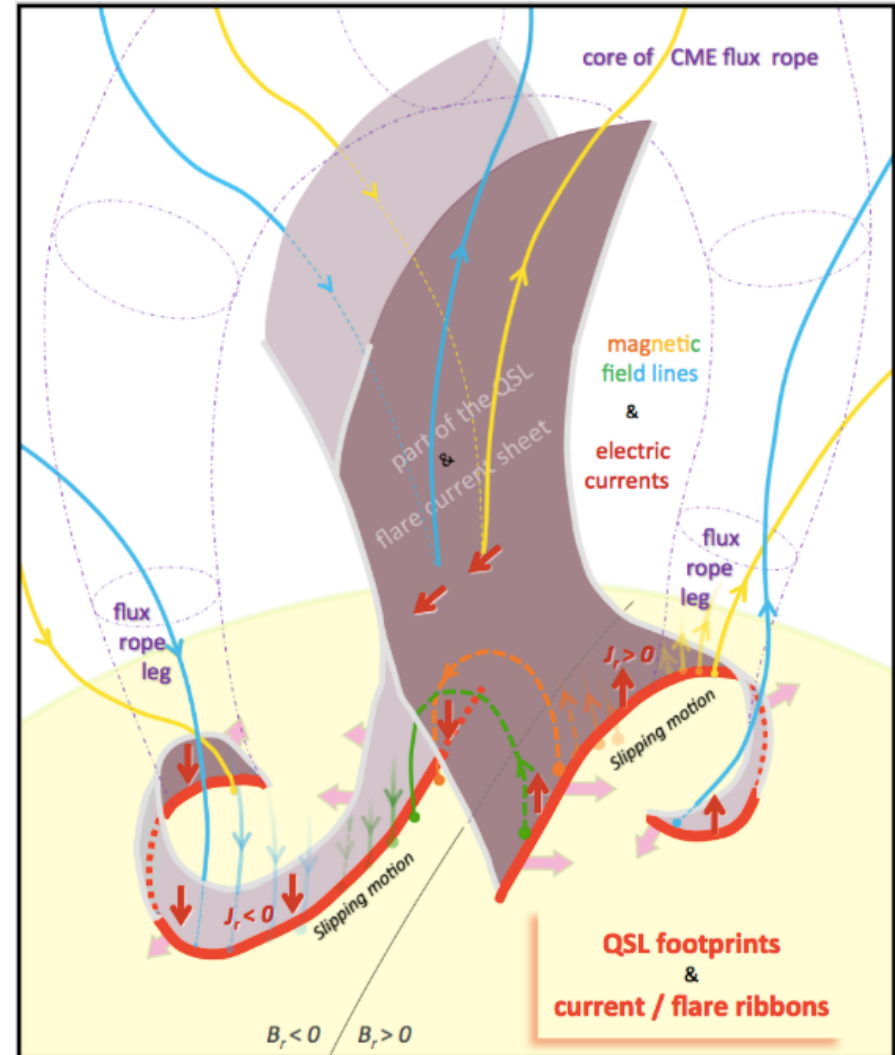
Observational history

- Liu, R et al. [2009ApJ...696..121L](#)
- Liu, R. & Wang, H. [2009ApJ...703L..23L](#)
- Liu, R. & Wang, H. [2009ApJ...703L..23L](#)
- Ji, H. et al. [2011ASInC...2..213J](#)
- Ji, H et al. [2011ASInC...2..213J](#)
- Sun, X. et al [2011sdmi.confE..63S](#)
- Liu, C. et al. [2012ApJ...745L...4L](#)
- Gosain, S. et al. [2012ApJ...749...85G](#)
- Shen, Y. et al. [2012ApJ...750...12S](#)
- Ji, H. et al. [2012EAS....55..229J](#)
- Simões et al. [2013ApJ...777..152S](#)
- ...
- Wang, J. et al. [2018ApJ...859...25W](#)

My new favorite cartoons

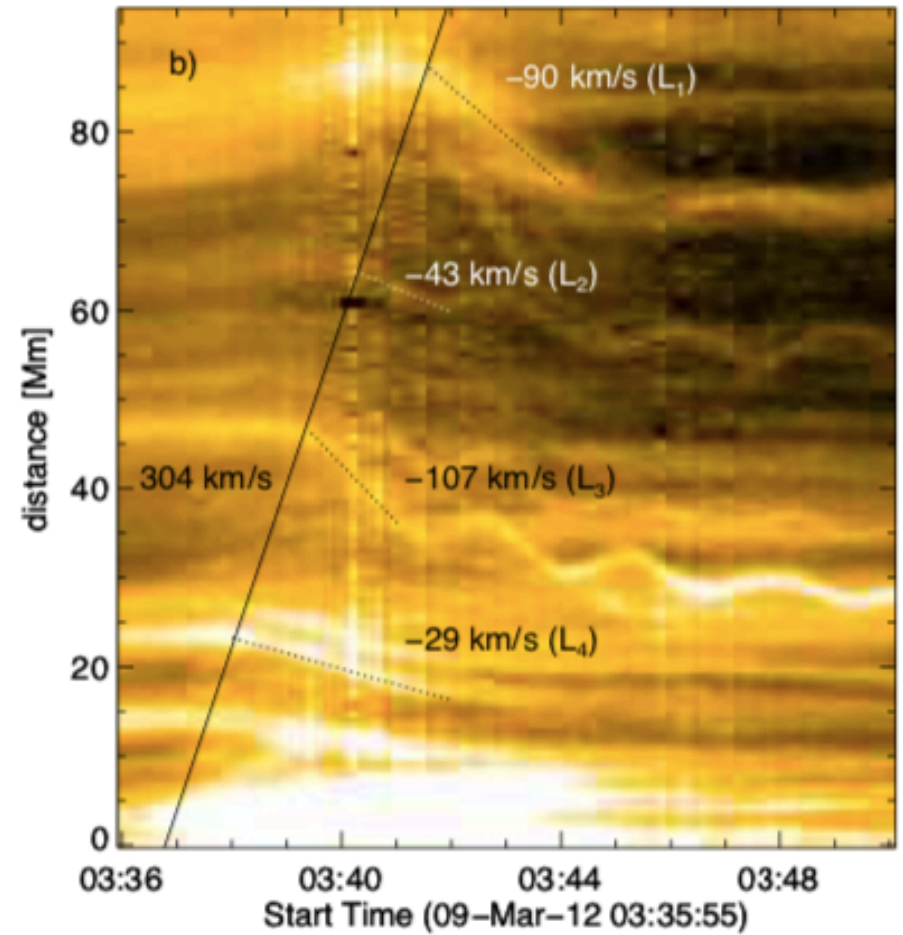
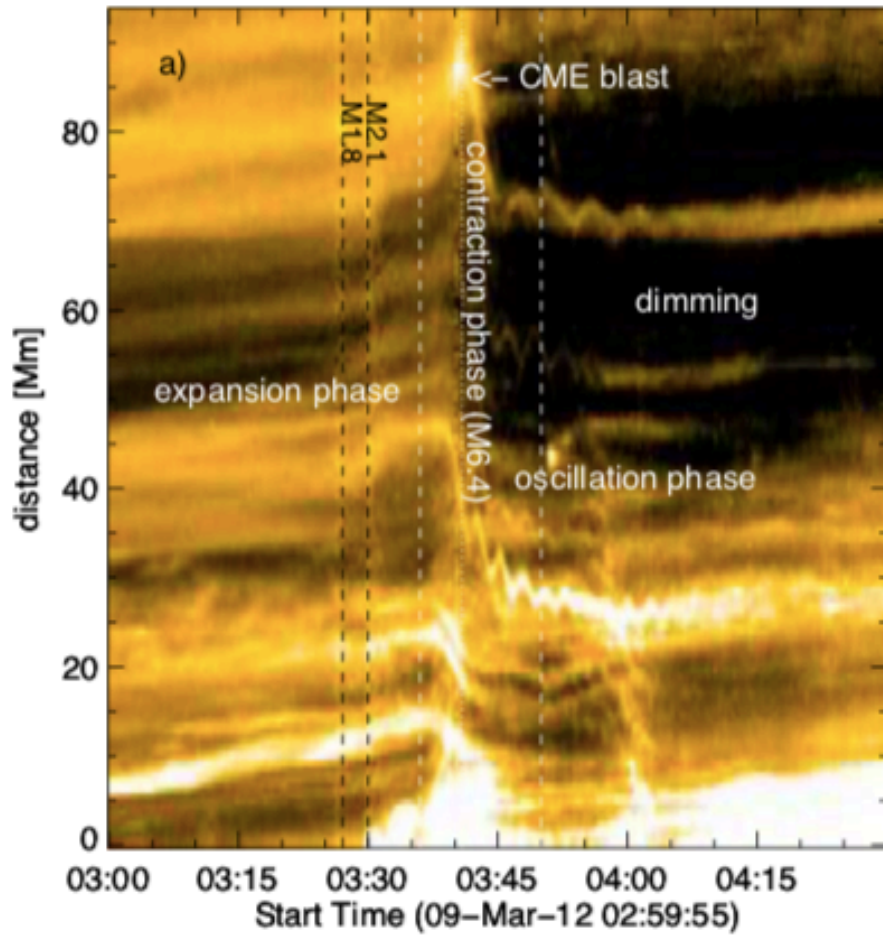


Russell et al. 2015



Janvier et al. 2013

Implosion, QPP, Dimming



Simões et al., 2013

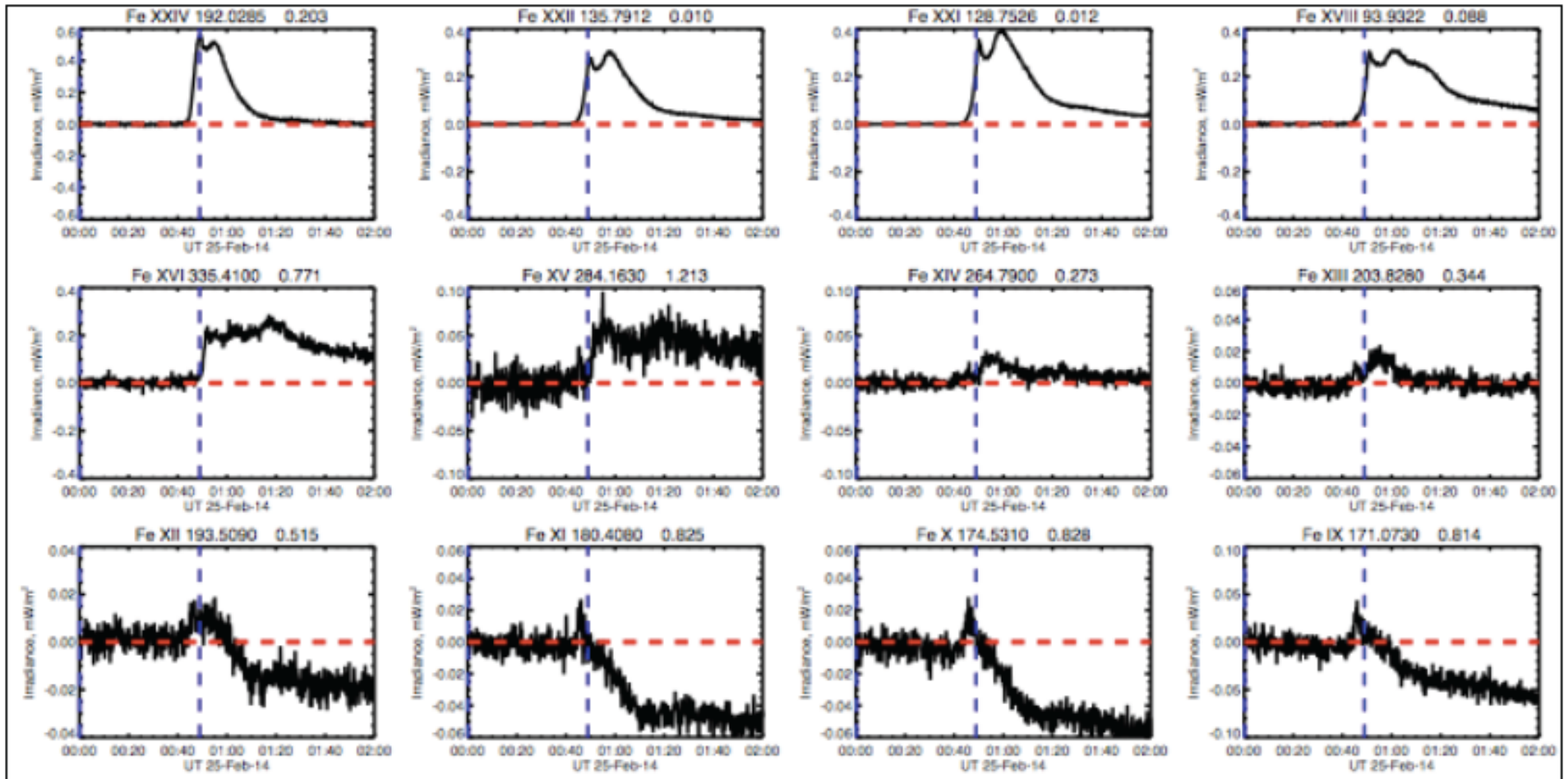
Conclusions

- A process of implosion, reducing $\int B^2 dV$ in the half-space, must supply the flare/CME energy on short time scales.
- Implosion is more fundamental than reconnection, and can be a driver.
- There are broad implications of the physical restructuring.
- Observationally it is tricky, probably because it is a process in 3 dimensions and we don't really know what the actual plasma instability is.

Challenges

- Observational: Can we detect implosions in non-eruptive flares?
- Theoretical: Does an implosion (ie, energy release) always require global magnetic reconnection?

SDO/EVE Spectroscopic Overview



An “Fe Cascade” plot, SOL2014-02-25

- Dimming (CME)
- Hot coronal sources
- Footpoint sources