Plasma Implosions in the Solar Atmosphere

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An eruptive flare **adds** free magnetic energy to the corona. How can this result from an instability?

Yohkoh and Shibata-sensei

- Yohkoh/SXT gave us our first comprehensive movie descriptions of flare plasma dynamics in its natural X-radiation
- During the decade of the 1990s, Shibata-sensei and I frequently met at ISAS, and I learned a great deal from him!

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from him! Mainly



of course!

The Hirayama model (1974)



Specific features **confirmed**:

- Filament channel
- Plasmoid
- Coronal overpressure
- Loop formation

Omissions TBD

- Plasma instability
- Particle acceleration

856 citations

Yohkoh-era flare cartoons





Shibata 1995 (451)

Moore 2001 (695)

Hudson 2000 (140)

https://www.astro.gla.ac.uk/cartoons/

Cartoon explanation



- This cartoon does not describe local properties ("reconnection"); it simply demonstrates conservation of energy
- Field lines shown in before/after steady states (solid, dotted)



Implosion, Flare, QPP, Dimming, Wave, HXR



Simões et al. 2013; Russell et al. 2015

• The Russell paper shows how all of these phenomena fit nicely together

Cartoon conclusions

- The implosion is fundamental, because it represents conservation of energy
- The Hirayama-Shibata-Moore (CSHKP) cartoon describes the later flare development
- The implosion cannot always be detected because it is in the (invisible) magnetic field
- Plasmoid eruption is a consequence, not a cause

Cartoon conclusion: The pressure reduction from the implosion drives the plasmoid



New (old) evidence for "tether-cutting"?

- Isothermal interpretations of GOES soft X-ray data invariably show "Hot Onset Precursor Events" (HOPEs) - only recently described
- The basic plasma evolution of a flaring loop does not predict this pattern
- The result is a pre-flare horizontal branch in (T,EM) diagnostic plots

The (EM, T) diagnostic diagram for a flaring loop



The (EM, T) diagnostic diagram for a flaring loop



Excellent recent example



- The HOPE (horizontal branch of [T, EM]) is almost always present
- Temperature tends to be fixed at 10-15 MK
- Emission measure rises roughly linearly
- HOPE appears before the impulsive phase

Why are hot onsets important for the global physics?

- The HOPE energy release is different from the standard flare pattern. We do not know yet whether it is steady-state or microflaring, but it is not "pre-heating" since dT/dt ~ 0
- Some AIA observations show that it may occur along the PIL that is going to flare, though not at the flare site necessarily
- A *Yohkoh* view of a single event (SOL2001-09-24) is in Fárník et al. 2003, and many other HOPE examples are elsewhere in the literature

Conclusions

- What we have learned about solar flares depends a great deal on Yohkoh and upon Shibata-sensei
- I think future work needs to focus on the lower atmosphere to understand "tether-cutting" physics and to understand how HOPE leads to eruption

If there is time, the "Slow HOPE" SOL2011-08-09T08:00





Message: The HOPE phase may be quite remote from the flare. It looks non-episodic here, but check out Fárník et al. 2003. How does this development lead to the major instability?